

Geometry and Strings 2018 - Ringberg

# The Refined Swampland Distance Conjecture in Calabi-Yau Moduli Spaces

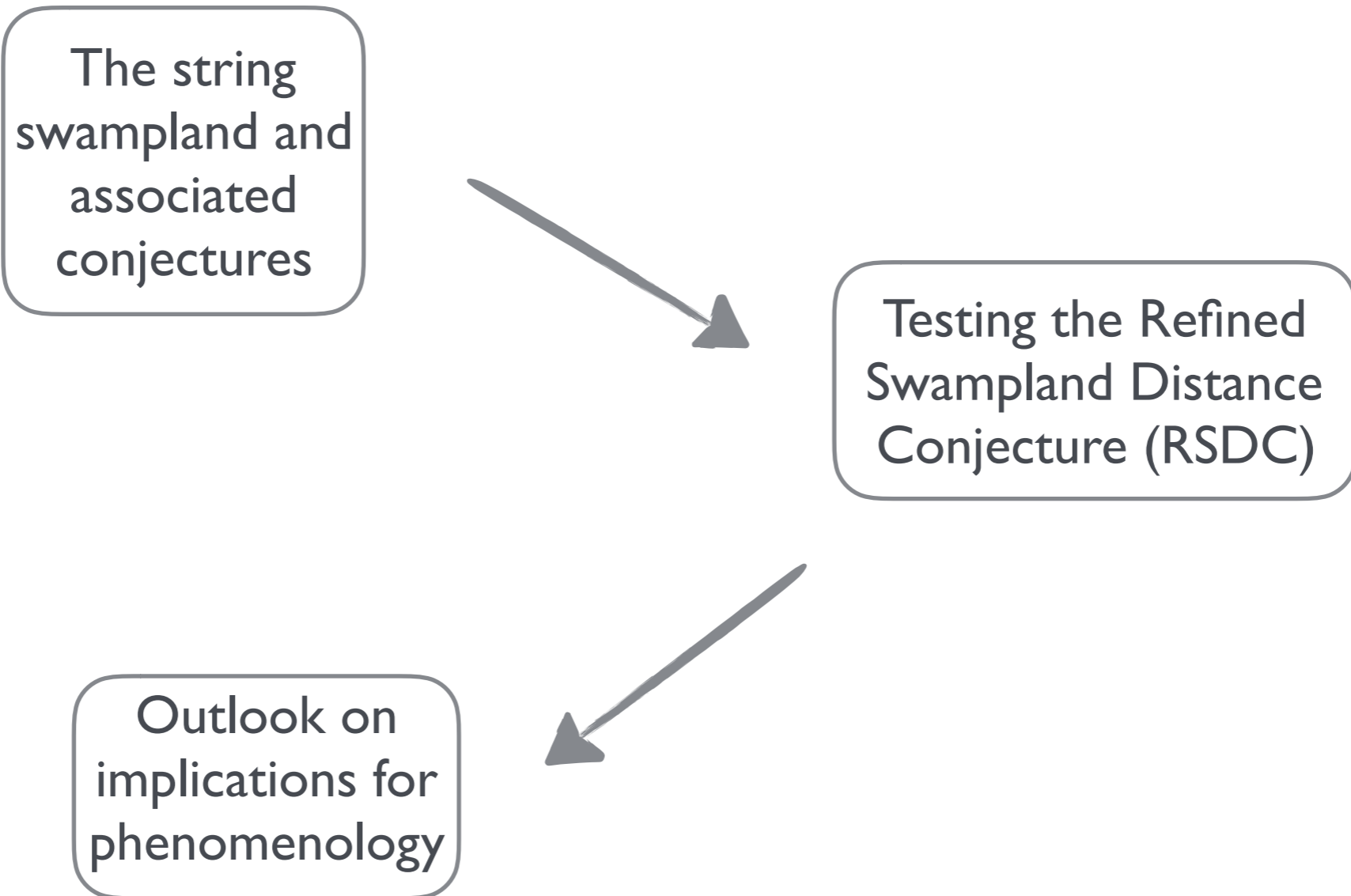
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# Outline



# The Swampland and Related Ideas

String theory has just celebrated its **50th birthday** 🎉 in Okinawa! String Phenomenology is a well-developed subject, addressing many problems in particle physics and cosmology from a top-down perspective.

**Many detailed constructions** have been developed to obtain: dS vacua, inflation, GUTs, etc...

Yet **general ideas about quantum gravity** and its realization in string theory appear to **challenge many of these models**.

The **(string) swampland** is the set of (seemingly consistent) effective field theories, which cannot be obtained from a consistent string construction.

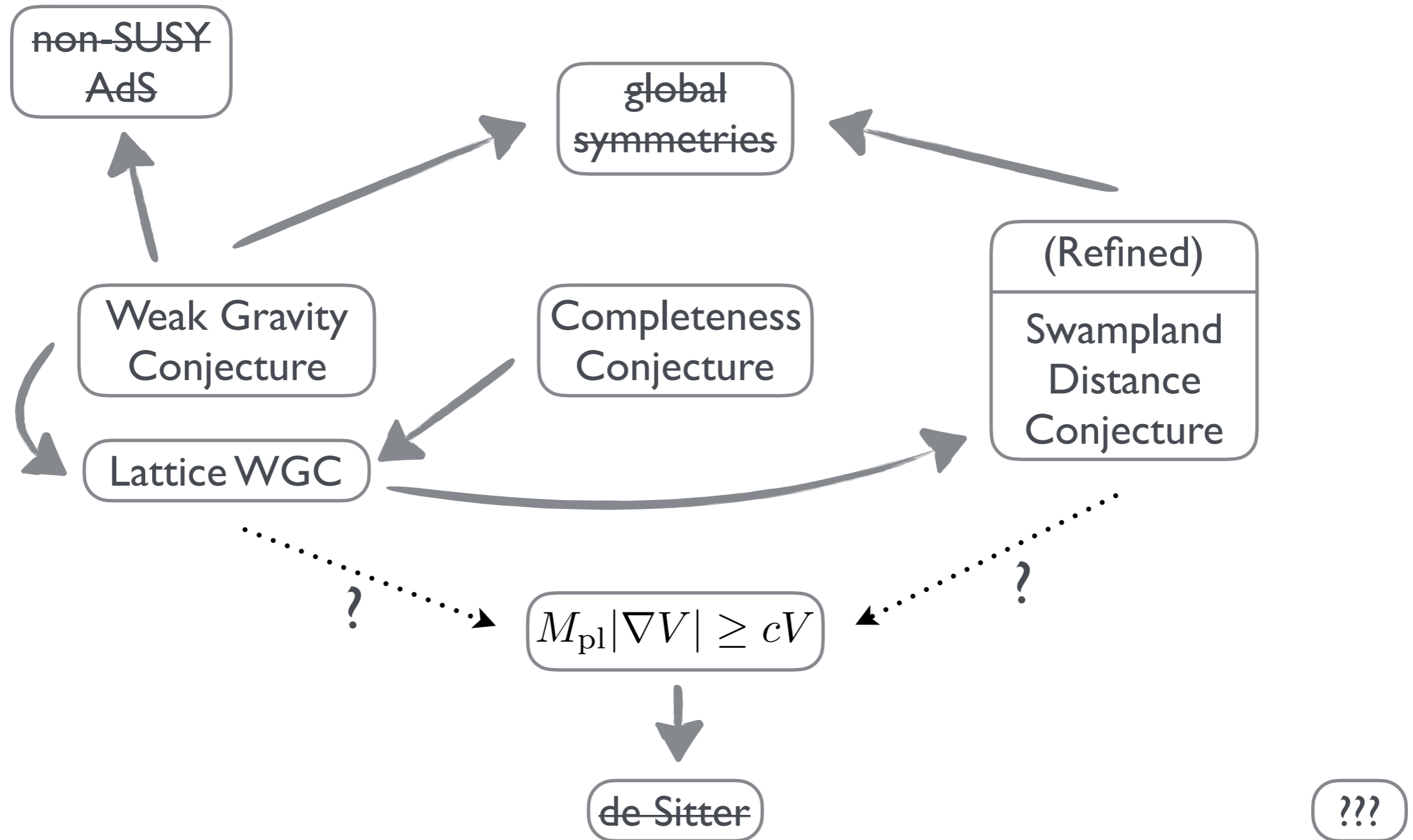
The **swampland as a blessing**: Knowing which field theories cannot be realized could actually lead to falsifiable predictions!!!

We want to **map out the boundary** of the swampland and explore the geometry of the landscape!



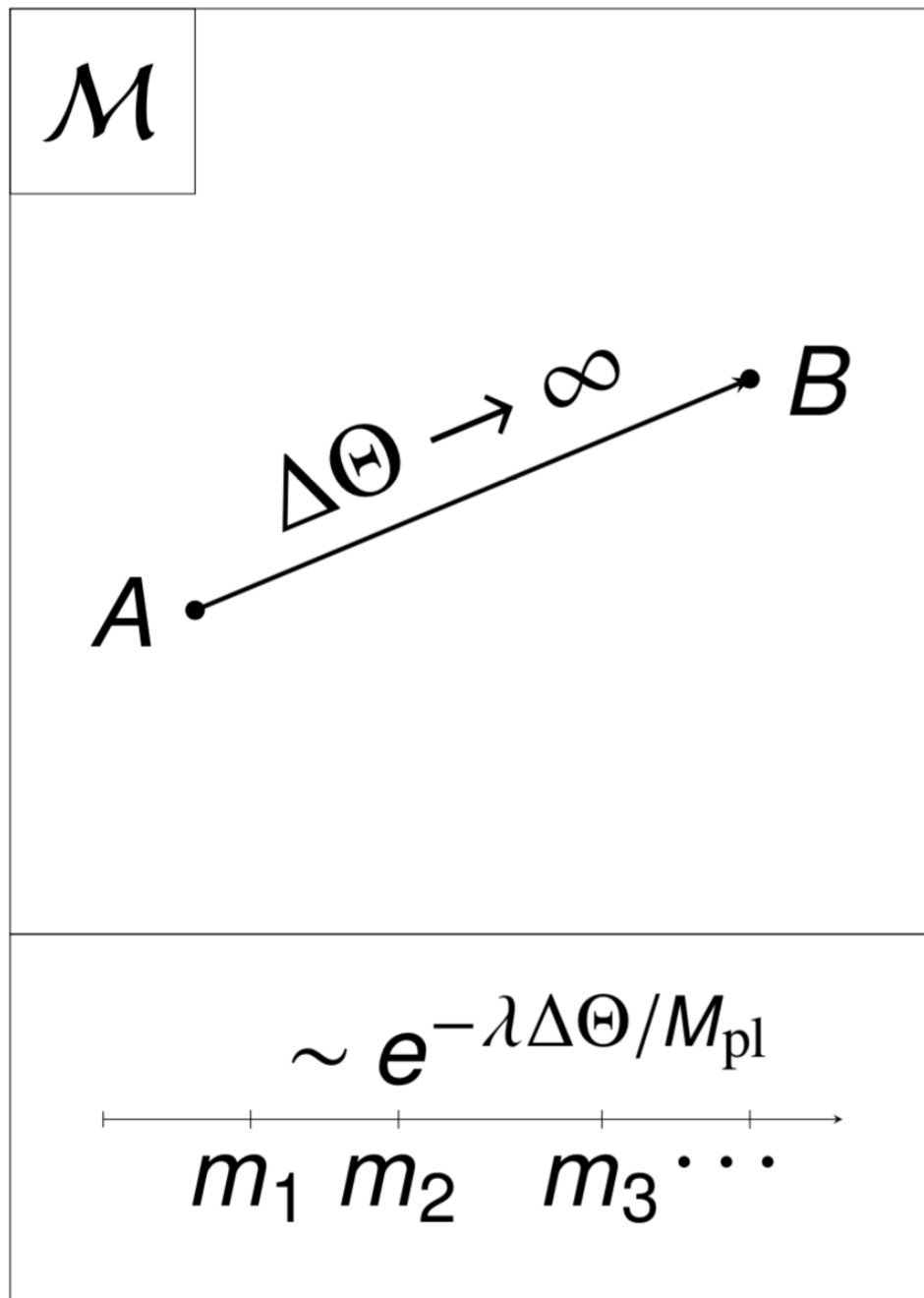


# A Web of Conjectures...



# The Swampland Distance Conjecture

[Ooguri, Vafa '06]



- Asymptotic displacements  $A \rightarrow B$  in continuous moduli space of quantum gravity
- Conjectured universal behavior of mass scale of an infinite tower of states

$$\Theta = \int_{\tau_A}^{\tau_B} d\tau \sqrt{G_{\alpha\beta} \frac{dx^\alpha}{d\tau} \frac{dx^\beta}{d\tau}}$$

- Casts doubt on validity of EFT for large field displacements

$$\Theta > \Theta_\lambda = \frac{1}{\lambda} = \mathcal{O}(1) M_{\text{pl}}$$

# Evidence

- Well known for **string theory on tori** (IIB on  $S^1$ ) [Ooguri, Vafa '06]

$$M_{\text{KK}} \sim \frac{1}{R^{\frac{8}{7}}} \quad M_{\text{W}} \sim R^{\frac{6}{7}} \quad G_{\text{RR}} \sim \frac{1}{R^2} \quad \Theta \sim \log \left( \frac{R_B}{R_A} \right)$$

- **Holds for  $N > 8$  supercharges** (moduli space is coset) [Cecotti '15]
- **Evidence also for  $N = 8$  supercharges** [Grimm, Palti, Valenzuela '18]  
[Blumenhagen, DK, Schlechter, Wolf '18]
- Evidence from **semi-classical arguments**, relating it to WGC: [DK, Palti '16]
- (Sub-)Lattice WGC predicts infinite tower of states with  $m \sim qgM_{\text{pl}}$
- In gravitational theory, scalar fields can grow at most logarithmically [Nicolis '08]

$$\Delta\phi < \frac{1}{\alpha} \log(r/r_*)$$

- Together with magnetic WGC bound on the energy density  $g(r) > \rho(r)^{\frac{1}{2}}$

Find that gauge coupling = mass drops at least exponentially in  $\Delta\phi$



# The Refined Swampland Distance Conjecture

[Baume, Palti '16; DK, Palti '16]

- SDC holds globally in simple moduli spaces (toroidal compactification)
- Generically expect the SDC to be **badly violated** at finite distances

Refined SDC quantifies this:

The universal exponential behavior sets in for

finite displacements  $\Theta_0$

of order the Planck scale or earlier

# Evidence

- Even **less evidence than for the SDC**
- The semi-classical argument gives a **hint**: Free Scalars can only support sub-Planckian variations. Inside sources  $\Delta\Theta > M_{\text{pl}}$  is indeed possible, but only logarithmic growth!
- Solid **evidence from string theory** has been **lacking**

→ **fill this gap!**

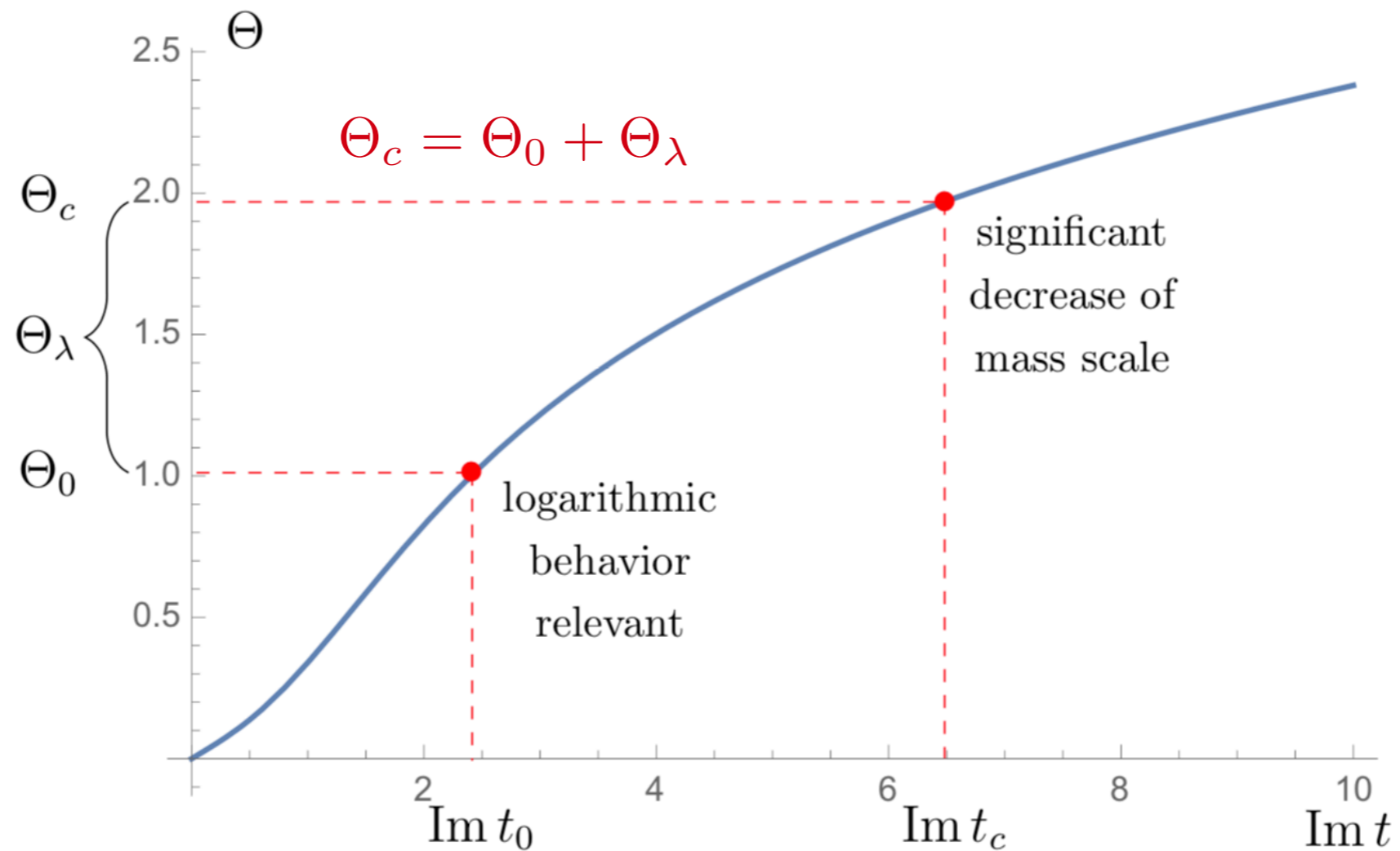
# Additional Evidence?

- The RSDC applies to moduli, i.e. flat directions. For Pheno, we really want it to apply it to fields with a potential (Inflaton,...).
- In fact, there is evidence that a similar mechanism is at work.
- (F-term) axion monodromy inflation: [Silverstein, Westphal '08; Marchesano, Shiu, Uranga '14]  
[Palti, Baume '16; Blumenhagen, Valenzuela, Wolf '17]
- Break axion shift symmetry by fluxes, but corrections to the effective potential controlled even in the trans-Planckian regime  $\Delta\Theta > M_{\text{pl}}$
- Axions do not control mass scales, should be safe from SDC
- For trans-Planckian axion, the axion valley moves into saxion direction (backreaction).  $s(\theta) = \lambda\theta$
- This implies the behavior predicted by the refined SDC

$$\Theta = \int K_{\theta\theta}^{1/2}(s)d\theta \sim \int \frac{d\theta}{s(\theta)} \sim \frac{1}{\lambda} \log(\theta)$$

# Objectives

- Test the Refined Swampland Distance Conjecture in CY moduli spaces

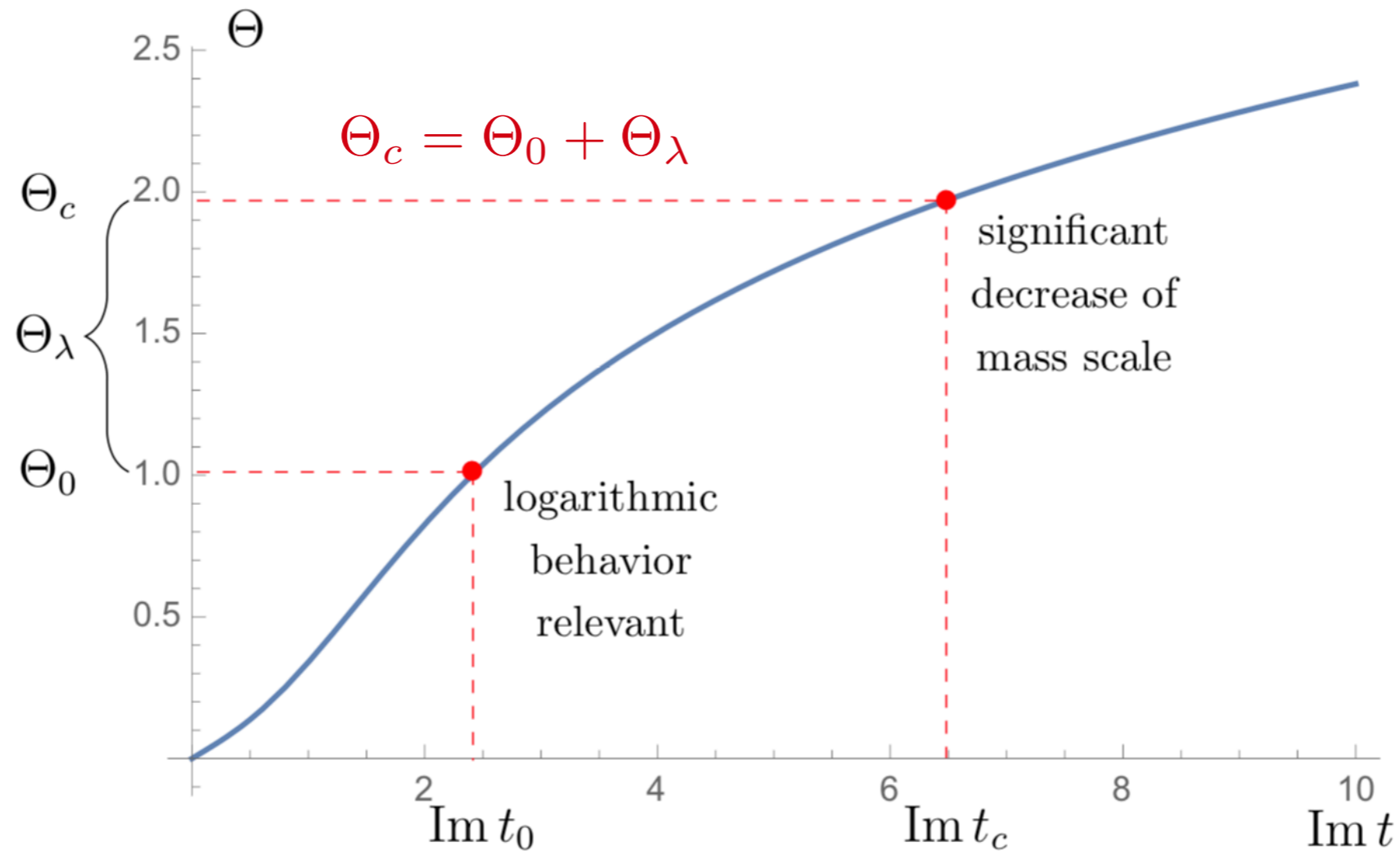


Prediction:

$$\Theta_0 < \mathcal{O}(1)M_{\text{pl}} \quad \Theta_\lambda < \mathcal{O}(1)M_{\text{pl}}$$

# Objectives

- Test the Refined Swampland Distance Conjecture in CY moduli spaces



Prediction:

$$\Theta_0 < \mathcal{O}(1)M_{\text{pl}} \quad \Theta_\lambda < \mathcal{O}(1)M_{\text{pl}}$$

# Calabi Yau Moduli Spaces

- IIA/IIB string theory on a Calabi-Yau M manifold with
 
$$h^{11} = \dim(H^{1,1})$$

$$h^{21} = \dim(H^{2,1})$$
- Low energy EFT: N=2 supergravity
- Moduli space of deformations of M splits into



IIA:  $h^{11}$  vector multiplets

$h^{21}$  hypermultiplets

IIB:  $h^{11}$  hypermultiplets

$h^{21}$  vector multiplets

- **Mirror symmetry:** duality between IIA on M and IIB on W (mirror CY)
  - Exchanges Kähler and CS moduli spaces

# Calabi Yau Moduli Spaces

- Metric on moduli space is determined by **Kähler potential**  $g_{\alpha\bar{\beta}} = \partial_{\alpha}\partial_{\bar{\beta}}\mathcal{K}$

$$\mathcal{K}_K = -\log\left(-\frac{i}{6}\kappa^{abc}(t_a - \bar{t}_a)(t_b - \bar{t}_b)(t_c - \bar{t}_c) + \xi + \mathcal{O}(e^{-2\pi it_a})\right)$$

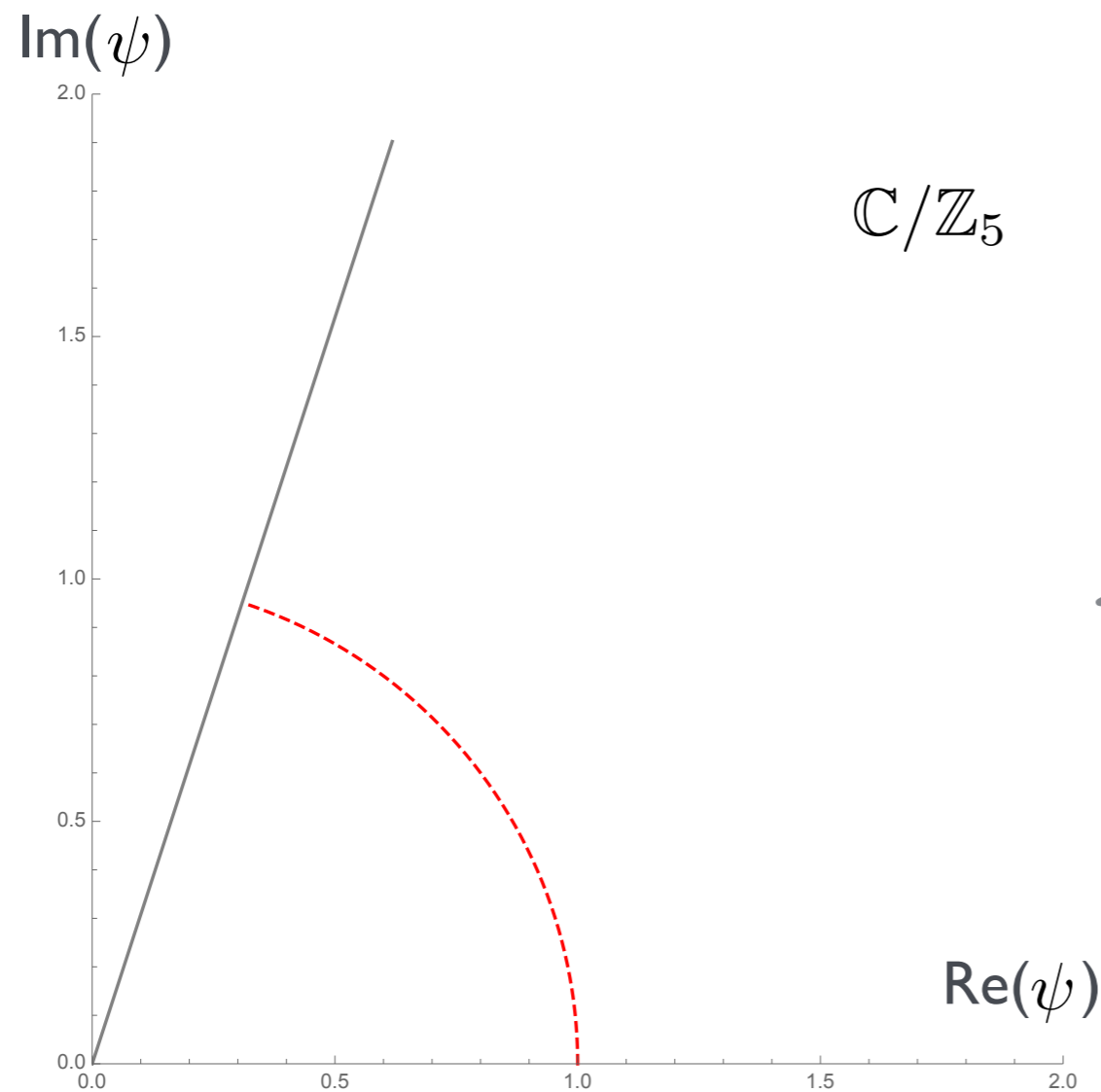
$$t_a = \int_{\Sigma_a} B + i \int_{\Sigma_a} J \quad a = 1, \dots, h^{1,1} \quad \text{compl. Kähler moduli}$$

$$\mathcal{K}_{CS} = -\log(-i\bar{\Pi}\Sigma\Pi) \quad \Pi_i(\Phi_{\alpha}) = \int_{A_i} \Omega(\Phi_{\alpha}) \quad i = 1, \dots, 2h^{2,1} + 2 \quad \text{periods}$$

- The **Kähler side** receives perturbative and non-perturbative **corrections**
- The classical result for the complex structure side is **exact**
- We focus on the Kähler side because of the obvious associated **tower of Kaluza-Klein states** (similar results apply for the CS sector)
- Use mirror symmetry as tool to compute the fully corrected Kähler potential and explore non-geometric regions of moduli space  $\text{Im}(t_j) = \mathcal{O}(1)$

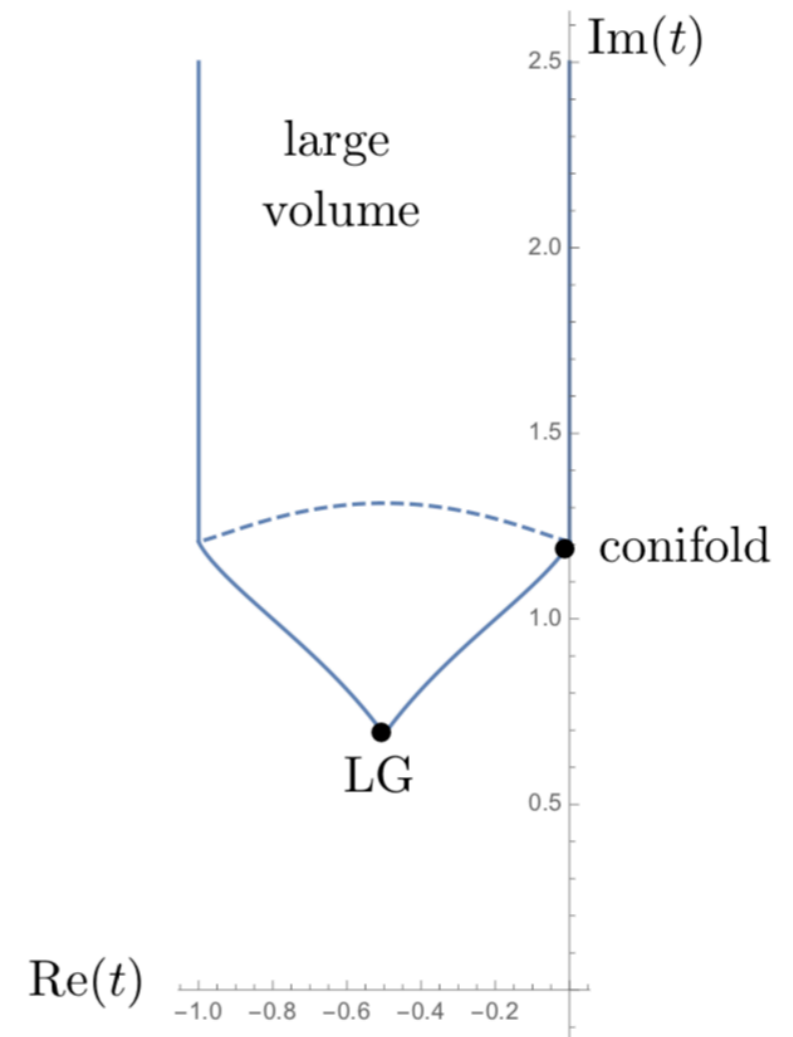
# CY moduli spaces and the RSDC

- **Example: (mirror) quintic**  $x_1^5 + x_2^5 + x_3^5 + x_4^5 + x_5^5 + 5\psi x_1 x_2 x_3 x_4 x_5 = 0$



$\mathbb{C}/\mathbb{Z}_5$

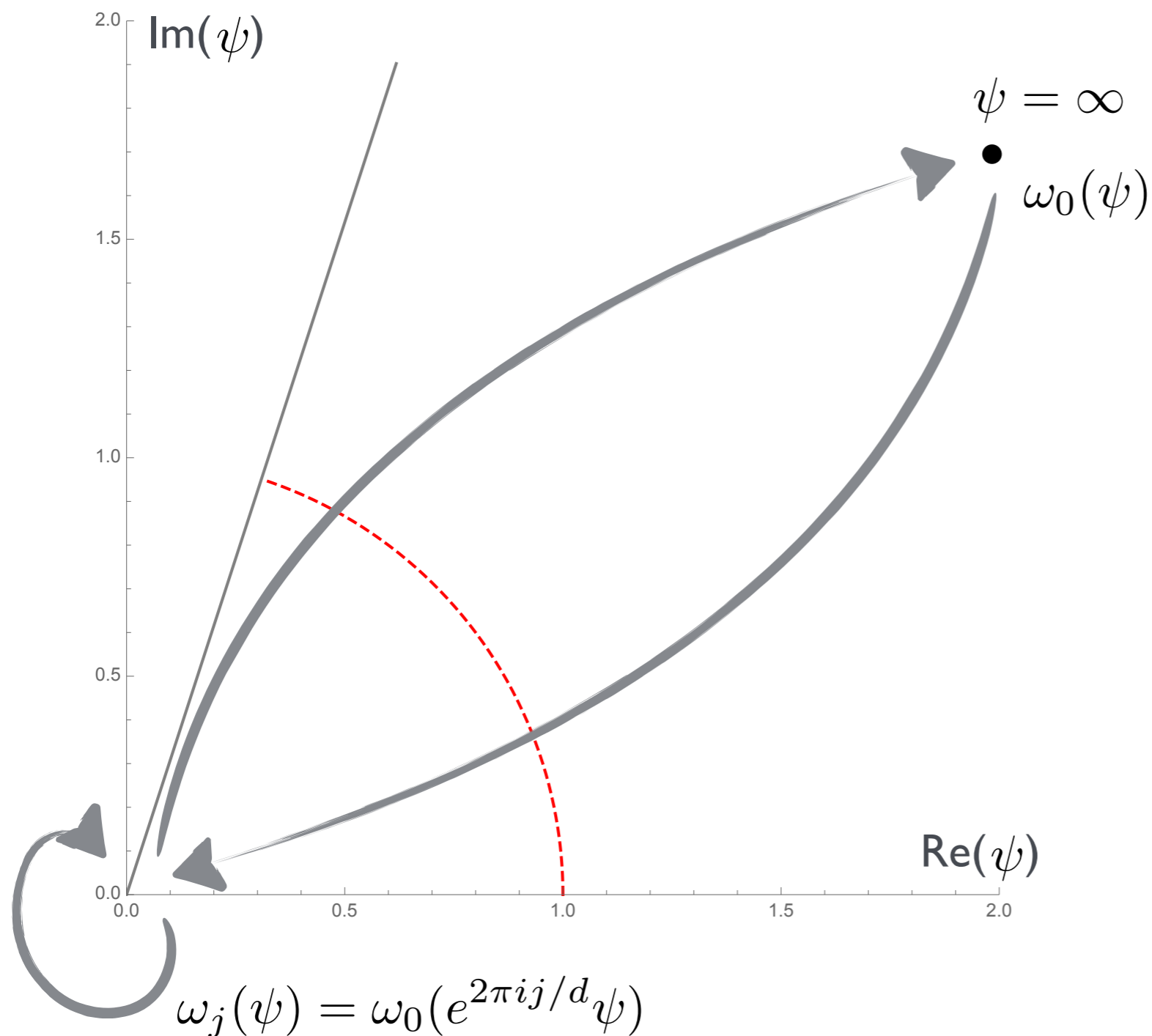
$\psi \rightarrow t(\psi)$   
mirror  
map





# Periods

- Well-known method to obtain Kähler potential on CS side and mirror map:



**Tedious, but can be done in a case by case analysis for  $h^{1,1}$  small**

[Berglund, Candelas, de la Ossa, Font, Hübsch, Jancic, Quevedo '93]

[Hosono, Klemm, Theisen, Yau '93]

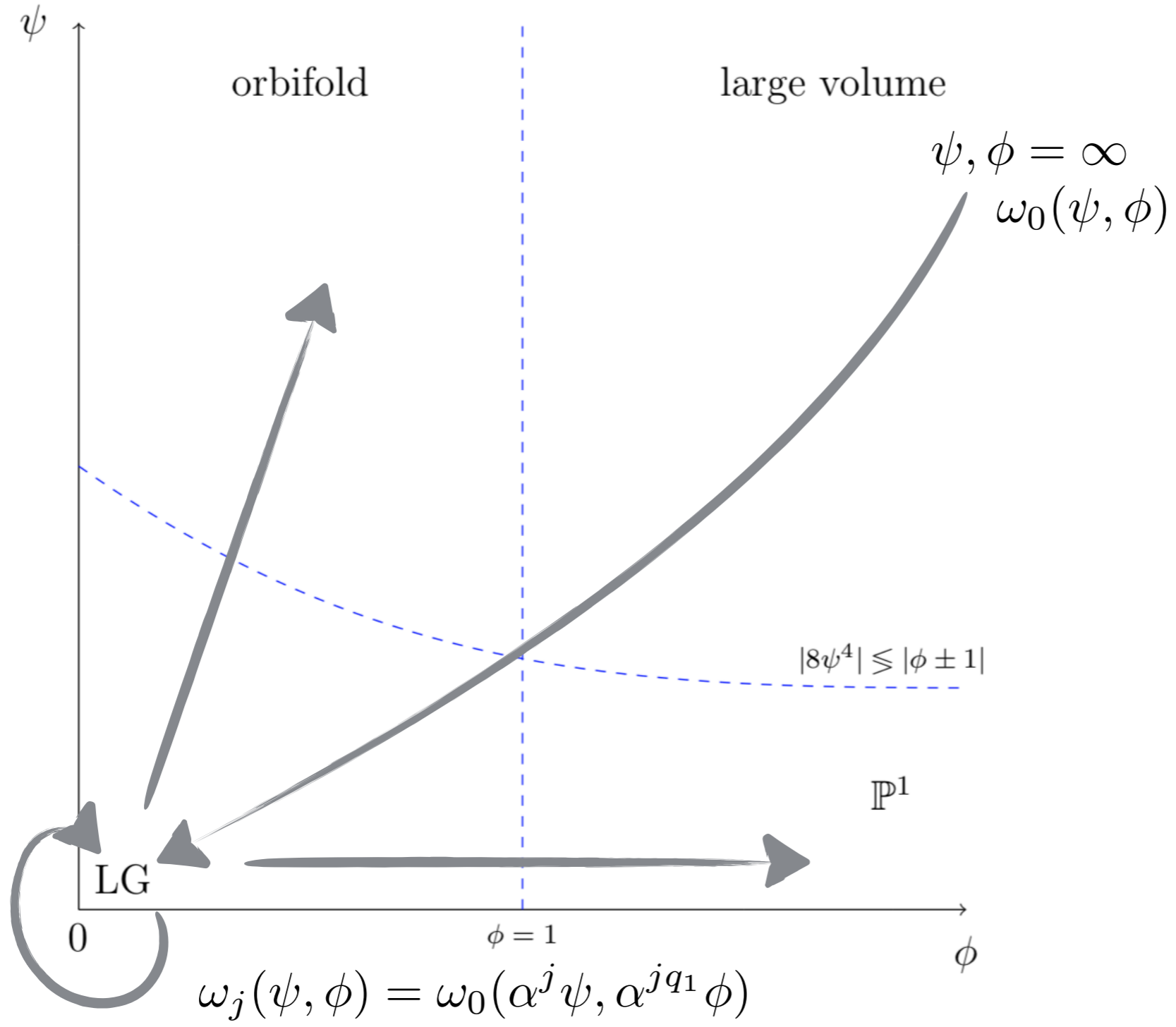
[Candelas, de la Ossa, Font, Katz, Morrison '94]

$$t = t(\psi)$$

$$\Pi = m\omega$$

# Periods for 2-dimensional moduli spaces

$$P = x_1^8 + x_2^8 + x_3^4 + x_4^4 + x_5^4 + \psi x_1 x_2 x_3 x_4 x_5 + \phi x_1^4 x_2^4$$



Analytic continuation is subtle, but periods can be written in terms of hypergeometric functions in different ways and standard techniques apply

# $\mathbb{P}^4_{11222}[8]$ and $\mathbb{P}^4_{11226}[12]$

[Berglund et al '94]

$$\underline{8|\psi|^4 > |\phi \pm 1|} \quad \omega_0(\psi, \phi) = \sum_{l=0}^{\infty} \frac{(q_1 l)! (d\psi)^{-q_1 l} (-1)^l}{l! \prod_{i=2}^5 \left(\frac{k_i}{d} (q_1 - q_i) l\right)!} U_l(\phi)$$

$$U_\nu(\phi) = \frac{e^{\frac{i\pi\nu}{2}} \Gamma\left(1 + \frac{\nu}{2}(k_2 - 1)\right)}{2\Gamma(-\nu)} \left[ 2i\phi \frac{\Gamma(1 - \nu/2)}{\Gamma\left(\frac{1+\nu k_2}{2}\right)} {}_2F_1\left(\frac{1 - \nu}{2}, \frac{1 - k_2\nu}{2}; \frac{3}{2}; \phi^2\right) + \frac{\Gamma(-\frac{\nu}{2})}{\Gamma\left(\frac{2+\nu k_2}{2}\right)} {}_2F_1\left(-\frac{\nu}{2}, -\frac{k_2\nu}{2}; \frac{1}{2}; \phi^2\right) \right]$$

$$\underline{8|\psi|^4 < |\phi \pm 1|} \quad \omega_0(\psi, \phi) = -\frac{2}{d} \sum_{n=1}^{\infty} \frac{\Gamma\left(\frac{2n}{d}\right) (-d\psi)^n U_{-\frac{2n}{d}}(\phi)}{\Gamma(n) \Gamma\left(1 - \frac{n}{d}(k_2 - 1)\right) \prod_{i=3}^5 \Gamma\left(1 - \frac{k_i n}{d}\right)}$$

**obtain all periods by**  $\omega_j(\psi, \phi) = \omega_0(\alpha^j \psi, \alpha^{jq_1} \phi)$ ,

# $\mathbb{P}_{11222}^4 [8]$ and $\mathbb{P}_{11226}^4 [12]$

Alternative representation:

$$\omega_j(\psi, \phi) = -\frac{2}{d} \sum_{r=1}^d (-1)^r e^{2\pi i j r / d} \eta_{j,r}(\psi, \phi)$$

$$\eta_{j,r}(\psi, \phi) = \frac{1}{2} \sum_{n=0}^{\infty} e^{i\pi n(j+1/2)} \frac{(2\phi)^n}{n!} V_{n,r}(\psi)$$

$$V_{n,r}(\psi) = N_{n,r} (d\psi)^r H_{n,r}(\psi)$$

$$H_{n,r}(\psi) = {}_{(d+1)}F_d \left( 1, \frac{n}{2} + \frac{r}{d}, \underbrace{1 + \frac{r}{d} - \frac{l_2 + 1 - \frac{n}{2}}{k_2}}_{i=3, \dots, 5}, \underbrace{1 + \frac{r}{d} - \frac{l_i + 1}{k_i}}_{l_i=0, \dots, k_i-1}; \underbrace{\frac{r+l}{d}}_{l=0, \dots, d-1}; \prod_{j=1}^5 k_j^{k_j} \psi^d \right)$$

Convenient for continuation to  $8|\psi|^4 > |\phi \pm 1|!$

# The Gauged Linear Sigma Model

- Can also compute directly on the Kähler side, using Witten's **gauged linear sigma model (GLSM)** description [Witten '93] [Jockers, Kumar, Lapan, Morrison, Romo '13]
- GLSM is N=(2,2) SUSY gauge theory in 2d. Varying the FI parameters leads to phase transitions, corresponding to phases of Kähler moduli space

- **Kähler potential is given by sphere partition function**  $e^{-\mathcal{K}} = Z_{S^2}$

$$Z_{S^2}(\xi, \bar{\xi}, Q, R) = \sum_{m_1 \in \mathbb{Z}} \cdots \sum_{m_s \in \mathbb{Z}} \int_{-i\infty}^{i\infty} da_1 \cdots \int_{-i\infty}^{i\infty} da_s Z_{\text{class}} Z_{\text{gauge}} Z_{\text{chiral}}$$

$$Z_{\text{chiral}} = \prod_{i=1}^M \frac{\Gamma\left(R_i/2 + \sum_{j=1}^s Q_{i,j} \cdot (a_j - m_j/2)\right)}{\Gamma\left(1 - R_i/2 - \sum_{j=1}^s Q_{i,j} \cdot (a_j + m_j/2)\right)}$$

$$Z_{\text{class}} = \prod_{j=1}^s e^{-4\pi i r_j a_j + i\theta_j m_j} \quad Z_{\text{gauge}} = 1$$

[Doroud, Gomis, Le Floch, Lee '13] [Benini, Cremonesi '15]

- Allows for **direct and algorithmic computation of the Kähler potential** without knowing the periods. Subtleties of analytic continuation are traded for subtleties in the evaluation of the integrals.

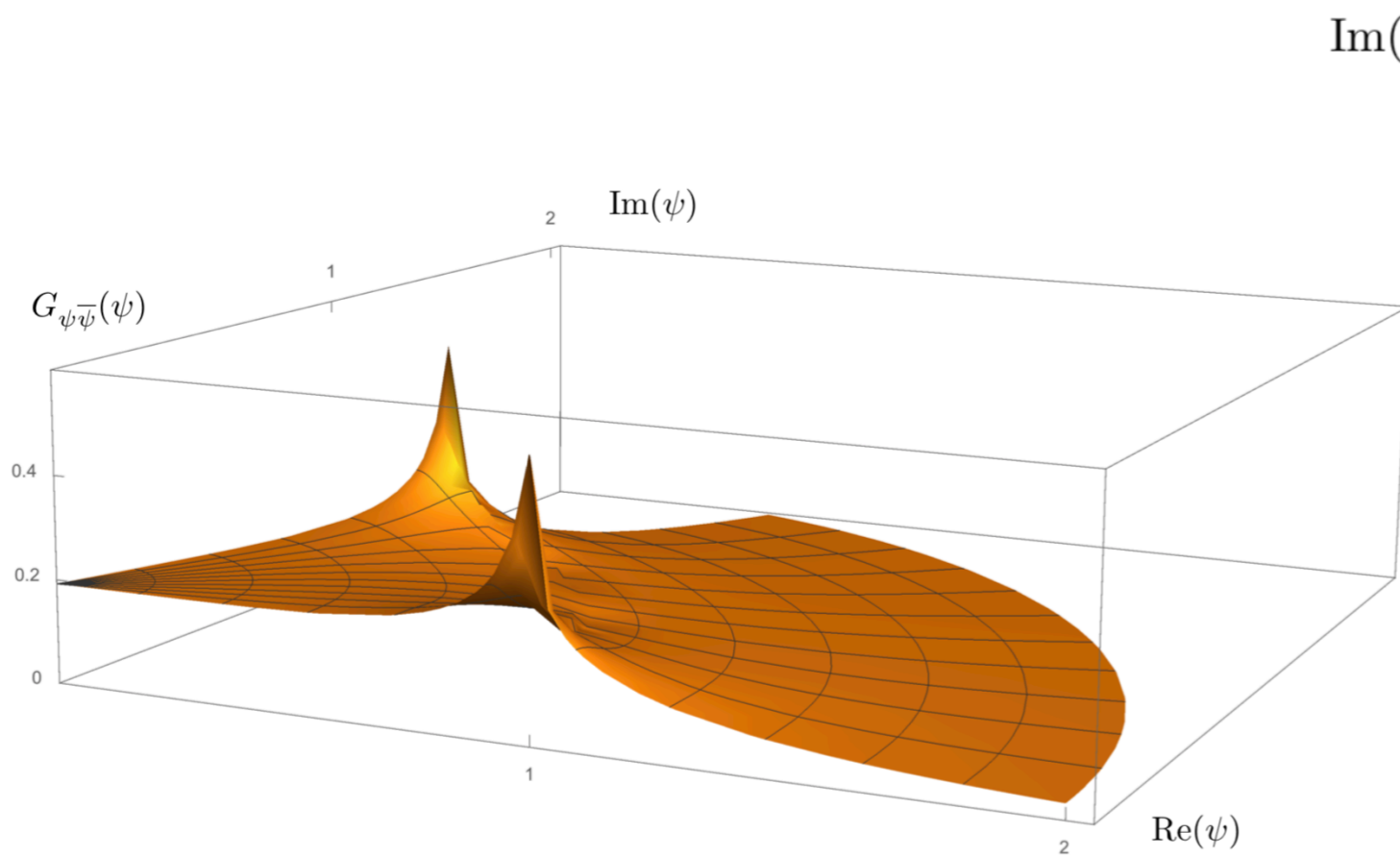
# The Quintic

$$x_1^5 + x_2^5 + x_3^5 + x_4^5 + x_5^5 + 5\psi x_1 x_2 x_3 x_4 x_5 = 0$$

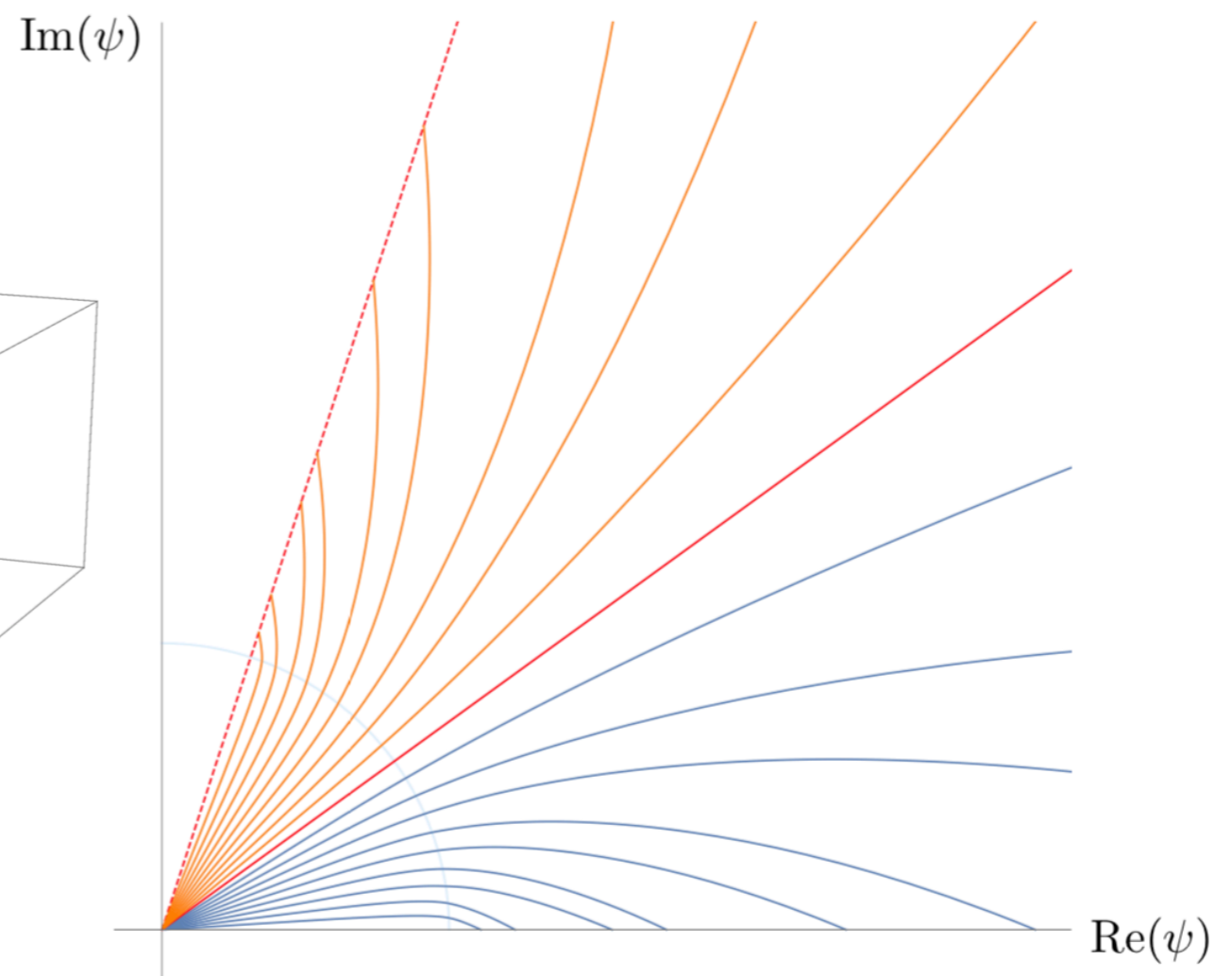
- **Necessary steps:**
- Compute metric, mirror map as described
- Determine the interesting regions in the moduli space (here: Landau-Ginzburg)
- Solve the geodesic equation numerically  $\frac{d^2 x^\mu}{d\tau^2} + \Gamma_{\alpha\beta}^\mu \frac{dx^\alpha}{d\tau} \frac{dx^\beta}{d\tau} = 0$
- Check consistency with the RSDC

# The Quintic

Metric



Geodesics



# Results

- Find distances **0.42-0.45** inside the LG phase
- $\Theta_\lambda$  varies because geodesics curve in axion direction

$$\Theta_0 \leq 0.45$$

$$\Theta_\lambda < 1$$

$\theta_{\text{init}} \cdot 60/\pi$	$\alpha_0$	$\alpha_1$	$\lambda^{-1}$	$\Theta_0$	$\Theta_c$
3	0.1315	0.2043	0.9605	0.4262	1.3866
4	0.1127	0.2099	0.9865	0.4261	1.4125
5	0.0998	0.2213	0.9780	0.4260	1.4040
6	0.0955	0.2294	0.9567	0.4259	1.3827
7	0.0818	0.2475	0.9611	0.4259	1.3869
8	0.0877	0.2592	0.9275	0.4258	1.3533
9	0.0808	0.2825	0.9253	0.4257	1.3510
10	0.0929	0.3093	0.8969	0.4257	1.3226
11	0.0998	0.3497	0.8845	0.4257	1.3102
12	0.1234	0.1662	0.8657	0.4256	1.2914

- Analyse all CYs with  $h^{1,1} = 1$  given by hypersurfaces in  $WCP$ , namely

$$\mathbb{P}^4_{111112}[6]$$

$$\mathbb{P}^4_{111114}[8]$$

$$\mathbb{P}^4_{111125}[10]$$

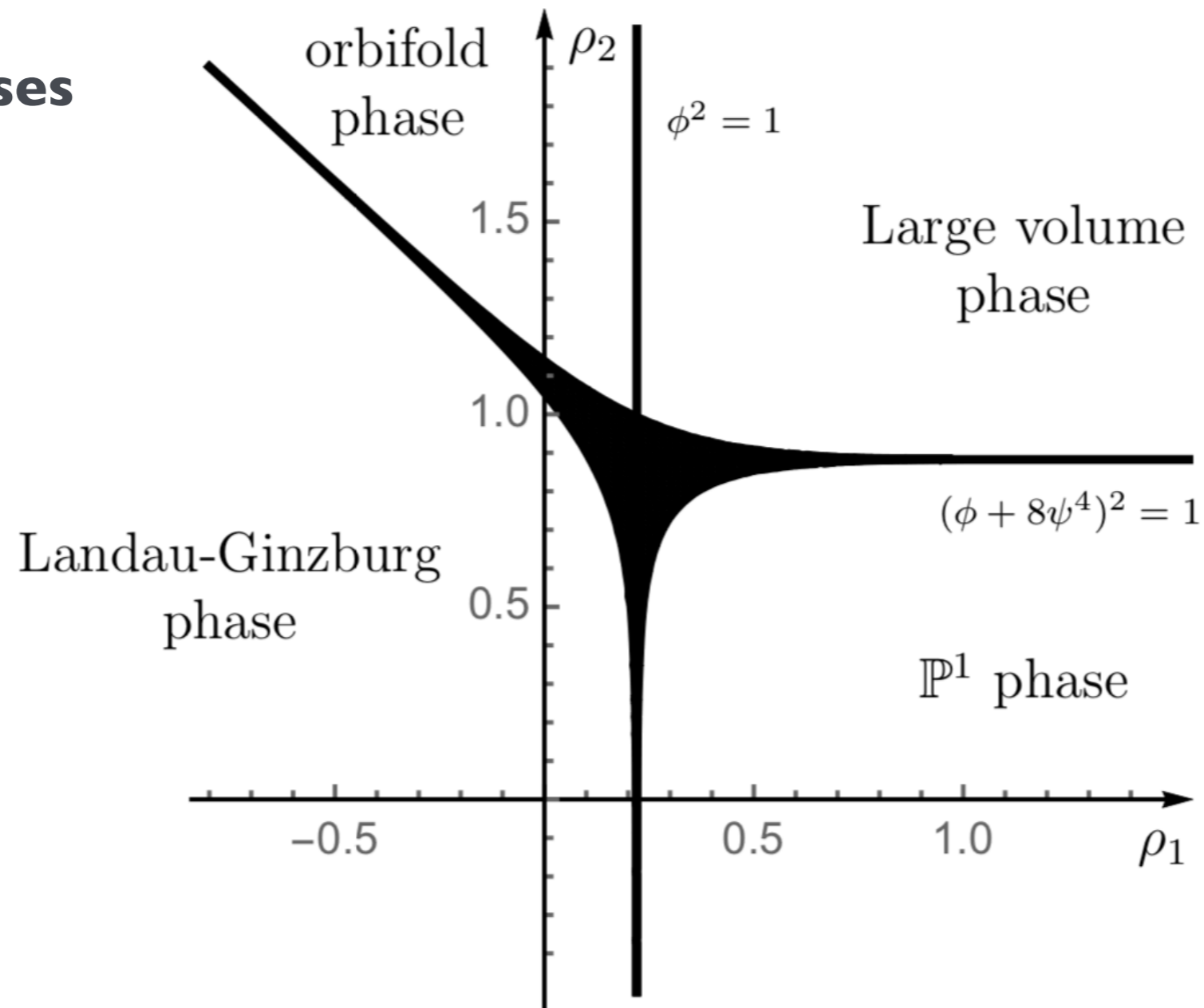
- All results in agreement with the RSDC, quintic is extremal

$$\Theta_c \equiv \Theta_0 + \Theta_\lambda \leq 1.4$$



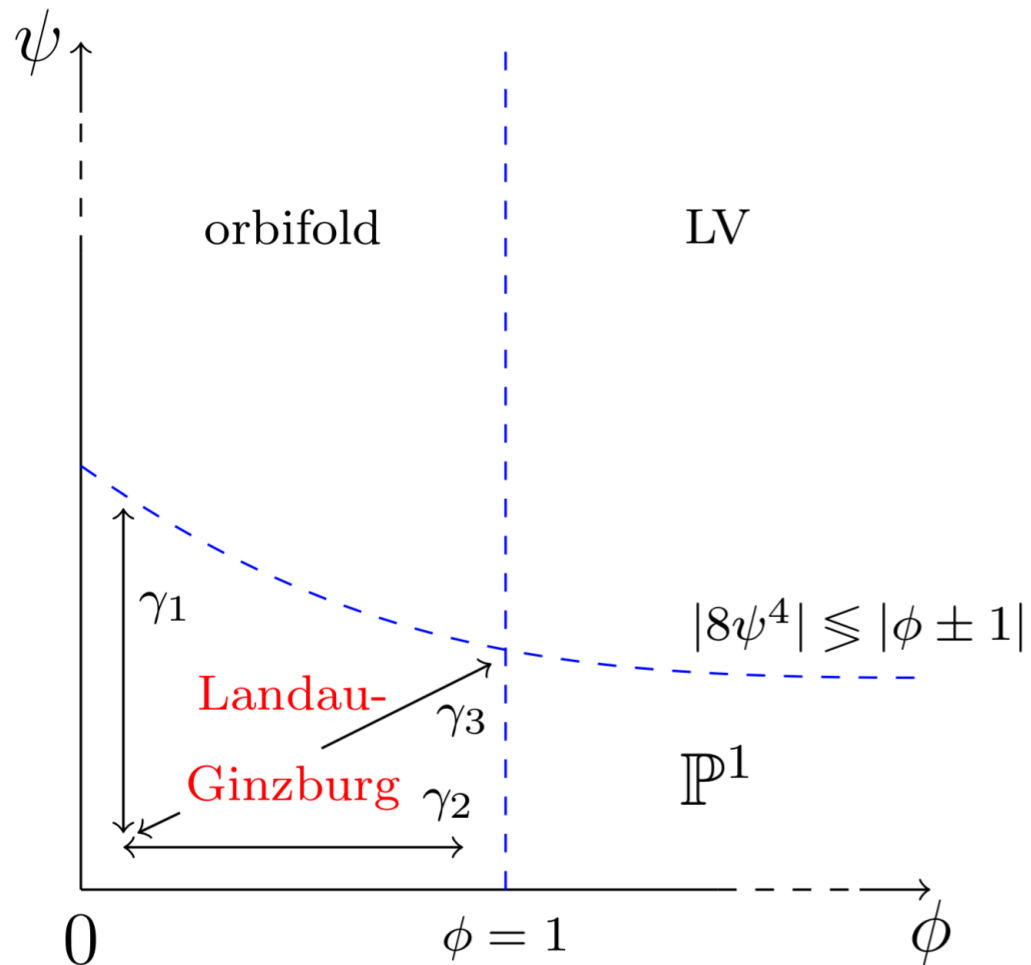
# CYs With $h^{1,1} = 2$

new feature:  
**hybrid phases**



[Aspinwall '94]

# $\mathbb{P}_{11222}^4 [8]: \text{LG phase}$



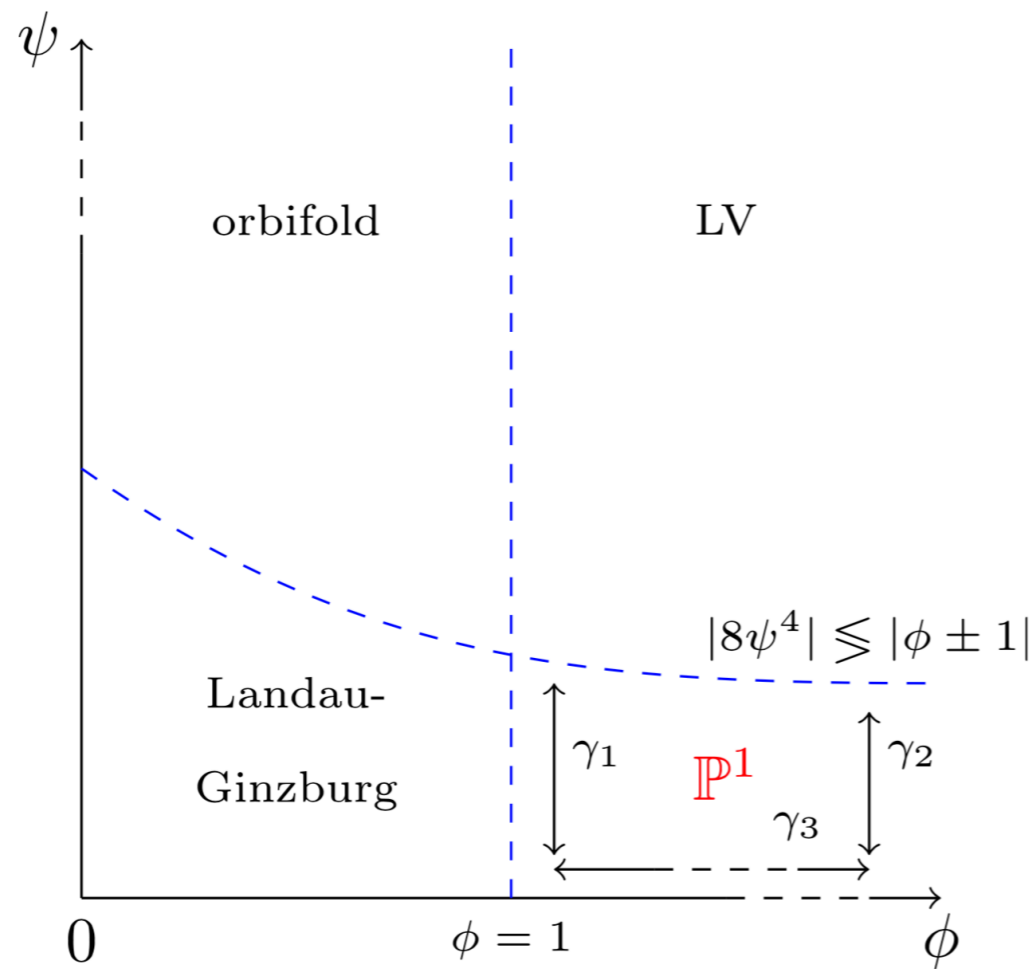
$$\Delta\Theta_1 = \int_{\gamma_1} d\psi \sqrt{G_{\psi\bar{\psi}}(\psi)} = 0.40,$$

$$\Delta\Theta_2 = \int_{\gamma_2} d\phi \sqrt{G_{\phi\bar{\phi}}(\phi)} = 0.24,$$

$$\Delta\Theta_3 = \int_{\gamma_3} d\tau \sqrt{G_{\mu\bar{\nu}}(\psi(\tau), \phi(\tau)) \frac{dx^\mu}{d\tau} \frac{d\bar{x}^{\bar{\nu}}}{d\tau}} = 0.36$$

Everything consistent with the RSDC!

# $\mathbb{P}_{11222}^4 [8]: \text{Hybrid Phase I}$



$$\Delta\Theta_1 = \int_{\gamma_1} d\psi \sqrt{G_{\psi\bar{\psi}}(\psi)} = 0.24$$

$$G_{\mathbb{P}^1}^{\text{asympt}} \simeq \begin{pmatrix} \frac{0.25}{|\phi|^2 (\log |\phi|)^2} & 0 \\ 0 & \frac{0.5905}{\sqrt{|\phi|}} \end{pmatrix}.$$

$$\begin{aligned} \Delta\Theta_2 &= \int_{\gamma_2} d\psi \sqrt{G_{\mathbb{P}^1, \psi\bar{\psi}}^{\text{asympt}}(\phi)} \\ &\simeq \sqrt{\frac{0.5905}{\sqrt{|\phi|}}} \cdot \sqrt[4]{\frac{|\phi|}{8}} = 0.46 \end{aligned}$$

Everything consistent with the RSDC!

# The Mirror Quintic: $h^{1,1} = 101$

$$P = x_1^5 + x_2^5 + x_3^5 + x_4^5 + x_5^5 + \psi x_1 x_2 x_3 x_4 x_5 + 100 \text{ other terms}$$

- Recent advances allow us to compute the **Kähler metric for the Landau-Ginzburg phase of the mirror quintic** [Aleshkin, Belavin '17]
- Computing geodesics in a 101 dimensional space numerically is hopeless
- Group deformations into equivalence classes under coordinate permutations  
 → left with 5 sets of deformations of cardinality (1, 20, 30, 30, 20)
- Compute proper lengths of collective displacements
- Compelling:  $\Theta_0 \sim \frac{1}{\#\text{fields}}$
- No parametric enhancement of  $\Theta_0$  in this way.

direction	$\Delta\Theta$
$\Phi_0$	0.4656
$\Phi_1$	0.0082
$\Phi_2$	0.0670
$\Phi_3$	0.0585
$\Phi_4$	0.0089

$$\frac{\Theta_0}{\text{phase}} \#(\text{phases}) \leq M_{\text{pl}} ?$$

# Implications for Cosmology

## Large Field Inflation

- Under pressure from several swampland conjectures
  - WGC constrains natural inflation
  - All models of large field inflation in tension with RSDC
  - OOSV  $|V'|/V > c = \mathcal{O}(1)$  puts pressure on slow roll [Obied, Ooguri, Spodyneiko, Vafa '18]

## Dark Energy

- If dS is in the swampland, what about quintessence?
- Borderline consistent with the OOSV conjecture, RSDC [Agrawal, Obied, Steinhardt, Vafa '18]

Are we missing something fundamental?

# Conclusion

- Refined Swampland Distance Conjecture passes many non-trivial tests in Calabi-Yau moduli spaces
- Diameter of non-geometric phases seems to approach zero as  $h^{1,1} \rightarrow \infty$
- Our analysis is case by case - it would be good to have a general argument!
- Many of the swampland conjectures turn out to be tightly related. Are there further relations?

**Thank You**